Performance of empirical models for diffuse fraction in Uruguay

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Abstract

Knowledge of diffuse solar radiation is required for the estimation of global irradiation on inclined surfaces or for estimating DNI for CSP applications. Since diffuse irradiance data is comparatively scarce relative to global horizontal irradiance (GHI) data, several methods are used to estimate the diffuse component of GHI. These methods have a local component and most of them have been developed using data recorded in the northern hemisphere, where long-term reliable measurements of diffuse irradiance are available. This work considers ten models for hourly diffuse irradiation and evaluates their performance, both in their original and locally adjusted versions, against data recorded at five sites from a subtropical-temperate zone in the southern part of South America (latitudes between 30° S and 35° S). The raw data has been quality-assessed by using a set of seven sequential filters which preserve the natural spread of the data while removing unphysical data points. The local adjustment and performance evaluation are done using random-sampling cross-validation techniques on an ensemble. The best estimates result from locally adjusted multiple-predictor models, some of which can estimate hourly diffuse fraction with uncertainty of 18% of the mean. *Keywords:* diffuse radiation, solar resource assessment, DNI

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Reywords. diffuse radiation, solar resource assessment, r

1. Introduction

The diffuse component (DHI) of the solar radiation reaching the ground is the result of several interactions between the incident solar (beam) radiation and the atmosphere. These processes can be described by physical models provided enough information on the current composition of the local atmosphere (i.e. aerosol type and density, water vapor column, Ozone column, among others) are available [1]. This detailed information is recorded at a few specialized ground measuring sites, such as those from Aeronet (http://aeronet.gsfc.nasa.gov/).

The separation of the beam and diffuse components of GHI is required before estimating direct normal irradiance (DNI) or global irradiance on inclined surfaces. Recent efforts in solar resource assessment in Uruguay have emphasized the characterization and modeling of GHI on several time scales [2–5], but there is little information available on diffuse radiation for

Since the final uncertainties in solar resource estimation correlate with financial risks in utility-scale projects, a reasonable knowledge of the uncertainties in each step of the calculations is important for the assessment of the performance of solar en-

this region. DHI is comparatively hard to measure accurately over long periods of time, so most available data sets include only GHI. A simple way to do this separation is to use phenomenological approaches, based on estimating DHI from a small set of easily measured or calculated predictor variables. These models refer to a definite time scale (typically an hour, a day or a month) and usually relate the diffuse fraction (the fraction of global horizontal irradiance (GHI) which is diffuse) to the clearness index and eventually other variables. They are not universal and several comparisons of their performance at different locations have been reported [6–11].

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ergy conversion technologies [12]. The uncertainty of a diffuse- ⁷⁰
fraction model will depend on the degree of climatic similarity ⁷¹
between the data sets used to develop the model and the cli- ⁷²
mate in which it is being evaluated. Localized assessments are ⁷³
necessary both to select the best model and to characterize its ⁷⁴
uncertainty. ⁷⁵

The diffuse fraction is not a function of clearness index alone. 76 38 Proposals with additional variables [8, 13–16] may have lower 77 39 uncertainties in diffuse fraction estimates at the expense of 78 40 higher complexity. Gueymard and Ruiz-Arias have recently 79 41 compared the performance of 140 diffuse fraction models pub-⁸⁰ 42 lished in the literature [6]. They used minute-based data from ⁸¹ 43 54 research-class stations distributed over four climatic regions 82 44 of the globe (only one of them is located less than 1000 km from 83 45 the area of interest in this paper) and characterized the regional 84 performance of each model. An important conclusion is that 85 47 no current separation model is truly "universal", in the sense to 86 48 have consistent accuracy over large climatic zones. In fact, the 87 49 diffuse fraction reflects the typical composition of the local at- 88 50 mosphere, which may be influenced by (natural or man-made) 89 51 phenomena affecting the water content or aerosol type and den- 90 52 sity at a specific region. Thus, diffuse fraction estimation is a 91 53 92 problem with an important local component. 54

Phenomenological separation models should be adjusted to 55 local data to remove most of their bias and significantly reduce 93 56 related uncertainties. However, these models are frequently 57 used as universal due to the absence of reliable local informa-58 tion on their performance. Many models for diffuse fraction 59 have been derived from DHI data taken at locations in the north-60 ern hemisphere, some of them at locations near densely pop-61 ulated areas, where these kind of measurements first became available. These models may not perform as well in locations 63 with different characteristics, as previously noted for Australia 64 by Boland [17]. 65

In this work, controlled-quality local diffuse irradiation data¹⁰² from five low-altitude sites with southern latitudes (between₁₀₃ 30° S and 35° S) is used to evaluate the performance of ten₁₀₄ well-known hourly diffuse-fraction models. A strong filtering₁₀₅ procedure is applied to the hourly data. For each model, both the original version and a locally adjusted version are evaluated against independent data using a standard cross-validation technique. Two frequently used models for daily and monthly average diffuse fraction are also evaluated and locally adjusted. Information is provided on the best way to estimate diffuse fraction for this and similar geographical regions on an hourly, daily and monthly basis. More importantly, the uncertainty associated to each estimation procedure is characterized, so that it may be accounted for in engineering calculations for solar energy projects.

The paper is organized as follows. In Section 2, the solar radiation database, the typical uncertainty for each site and the filters applied on the raw data are discussed. In Section 3, hourly diffuse fraction models are briefly described and evaluated against local data. In Section 4, all hourly models are adjusted to local data and re-evaluated on a per-site basis using several common statistical indicators. A global adjusted version of each model is defined and evaluated. In Section 5, the data is reduced to daily totals and two daily and monthly average models for diffuse fraction are implemented, locally adjusted and evaluated. Finally, In Section 6 our conclusions are summarized.

2. Ground data

The data used in this work consists of simultaneous data sets for hourly global and diffuse horizontal irradiation from five sites located in a sub-tropical temperate zone of the southeastern part of South America with homogeneous climatic characteristics shown in Figure 1. The area has a marked seasonality, no significant volcanic activity, low population density (except for the Buenos Aires metropolitan area) and it is not heavily industrialized.

2.1. Description of data sets

The location, instruments and number of hourly records (simultaneous global and diffuse irradiance) for each site are listed in Table 1. All sites are at low altitudes, with the highest (AR)

		Location				Time period	Ins			
Site	Code	LAT (°)	LON (°)	ALT (m)	owner	start - end	GHI	DHI	hours	uncertainty
Montevideo	AZ	-34.92	-56.17	58	LES	03/2014 - 08/2013	Delta-T SPN1 [4%]	Delta-T SPN1 [7%]	7961	9%
Salto	SM	-31.27	-57.89	41	INUMET	06/1998 - 12/2003	KZ CM11 [3%]	KZ CM11 + s-ring [4%]	20594	6%
Luján	LU	-34.58	-59.05	20	GERSolar	01/2011 - 06/2012	KZ CMP11 [3%]	Eppley 8-48 + s-ball [5%]	5934	7%
Artigas	AR	-30.40	-56.51	136	LES	02/2014 - 12/2015	Delta-T SPN1 [4%]	Delta-T SPN1 [7%]	7613	9%
Treinta y Tres	TT	-33.27	-54.17	20	LES	02/2014 - 12/2015	Delta-T SPN1 [4%]	Delta-T SPN1 [7%]	6634	9%

Table 1: Location of the measurement sites considered in this work (see Fig. 1 for the geographical distribution of the sites). Time period (month/year) and pyranometer manufacturer, model and estimated uncertainty for hourly averages. The method used for DHI measurement at SM was a Kipp & Zonen (KZ) CM-121 shadow-ring (s-ring). At LU it was a shading ball assembly (s-ball) based on a SOLYS2 tracking system. The last column indicates the valid daytime hours (F0 level, see Table 2) with simultaneous GHI and DHI measurements. And the last column is the overall estimated uncertainty for the normalized data from each site.



Figure 1: Location of the measuring stations considered in this work. Other details are provided in Table 1.

at 136 m above sea level. Except for AZ, all sites had a daily₁₂₁
cleaning routine by local staff. For AZ, the cleaning routine₁₂₂
was performed on a weekly basis. No ventilation devices where₁₂₃
used.

The AZ site is located at the roof-top of the School of Engi-125
neering at Montevideo, an urban coastal location. GHI and DHI126
were measured and recorded at one-minute intervals between127
2011 and 2013 using a new Delta-T SPN1 pyranometer. The128
data was recorded at 1-minute intervals using a Fisher-Scientific129
DT80 datalogger connected by cable to the internal network. 130
The SM site was at a supervised meteorological station run131

by the National Meteorological Service (INUMET), located₁₃₂

close to the Salto air field in a semi-rural location. GHI and DHI were measured and recorded at 15 minute intervals, during six years using two CM11 (Secondary Standard) Kipp & Zonen (KZ) pyranometers. The raw data was recorded with a Campbell Scientific CR1000 datalogger and it was provided to us without any processing. DHI was measured with a manually adjusted shadow-band, also from KZ. The DHI data was corrected using the isotropic correction factor [18] as provided by the band manufacturer, $f = (1 - S)^{-1}$, with

$$S = \frac{2\theta_0}{\pi} \cos \delta \left(\omega_s \sin \phi \sin \delta + \sin \omega_s \cos \phi \cos \delta \right)$$
(1)

where $\theta_0 = 0.185$ rad is the view angle of the shadow ring, δ the solar declination angle, ω_s the sunset hour angle (in radians) and ϕ the site latitude. This factor, which at the relevant latitudes varies yearly between 1.05 and 1.14, accounts for the portion of hemispherical sky radiation blocked by the shadow band under the assumption of an isotropic distribution of diffuse irradiance. According to the manufacturer, the correction from Eq. (1) is accurate to $\pm 0.5\%$. A comparison of several correction methods for diffuse irradiance measurements based on shadow rings [19], suggests a typical uncertainty of 4% with respect to a shading-ball assembly measurement, provided secondary-standard class pyranometers are used in both cases. Based on these considerations and on our own verifications, we estimate a typical uncertainty of 3% for GHI and 4 % for DHI hourly data from this site, relative to the overall mean values.

The LU site is at a specialized research laboratory (GERSO-

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LAR) of the National University at Luján (Argentina) located₁₇₁ 133 in a semi-rural area 50 km from the city of Buenos Aires. Three172 134 independent measurements (GHI, DHI and DNI) were recorded 135 at 1-minute intervals between 2011 and 2012. Global irradiance 136 was measured with a KZ CMP11 pyranometer, diffuse irradi-137 ance with a Black and White Eppley 8-48 pyranometer using a 138 shade-ball assembly and the beam component was measured 139 with an Eppley NIP pyrheliometer. These instruments were 140 mounted on a new SOLYS2 tracking system from KZ. The cal-141 ibration of pyrheliometers and pyranometers was done by peri-142 odic comparisons against a Kendall Absolute Cavity radiometer 143 used as secondary standard. Further details on this data set can 144 be found in Ref. [20]. The estimated uncertainty for data from 145 this site is 3 % and 5 % for hourly measurements of GHI and 146 DHI respectively, relative to the overall mean values. Although 147 the shading-ball assembly method for diffuse measurements is 148 potentially more accurate than a shadow-ring measurement, the 149 Eppley 8-48 pyranometer used for diffuse measurements has a 150 typical uncertainty of 5%, as declared by the manufacturer. 151

The TT site is part of an experimental station managed by 152 the National Institute of Agronomical Studies (INIA) located¹⁷³ 153 in a rural area, about 50 km from the nearest populated areas.¹⁷⁴ 154 The AR site is at a meteorological station run by the National¹⁷⁵ 155 Meteorological Service (INUMET) located in a semi-rural area,176 156 15 km from the town of Artigas. Two new Delta-T SPN1 pyra-177 157 nometers were installed by our laboratory at these sites in febru-178 158 ary 2014, and data was recorded at 1-minute intervals and sent¹⁷⁹ 159 on a daily basis to a dedicated server via the cellular (GSM) net-180 160 work. Data from these sites recorded between 2014 and 2015181 161 has been used in this work. At both sites, redundant GHI mea-182 162 surement using KZ CMP11 and CMP6 pyranometers were in-183 163 stalled and all the instruments received daily maintenance from184 164 the local staff. The sites AZ, TT and AR are part of a contin-185 165 uous solar radiation measurement network maintained by our₁₈₆ 166 laboratory since 2010. The pyranometers in this network are187 167 calibrated at our laboratory at two-year intervals, following ISO₁₈₈ 168 9847:1992(E) norm procedures [21]. The Secondary Standard189 169 used as a reference is a KZ CMP22 pyranometer calibrated₁₉₀ 170

against the World Radiometric Reference (WRR) at the World Radiation Center (WRC) at Davos in August 2014.



Figure 2: Hourly diffuse fraction, f_d , vs. clearness index, k_t , for all sites filtered to F7-level are shown in black. Unfiltered (F0-level) data for all sites is shown in the background (gray).

The SPN1 pyranometer has no moving parts and can operate over long periods of time without human intervention other than the cleaning procedures required by all hemispherical instruments. These radiometers are robust and allow continuous measurement of global and diffuse irradiance at remote locations in a cost-effective way as compared to other alternatives, such as rotating shadowband radiometers or tracker-based measurements. DHI data from SPN1 instruments account for about 46% of the data used in this work (Table 1) so we shall briefly discuss the accuracy of these instruments.

This pyranometer uses an array of seven thermopile sensors, each of them calibrated to consistently measure solar irradiance. A special mask under its dome shades at least one of the sensors from direct sunlight while leaving at least one of them exposed to direct sunlight, at all times and locations. This mask blocks approximately half of the hemisphere and the instrument computes individual values for GHI and DHI using a simple algorithm based on the maximum and minimum irradi-

ance it measures in its seven sensors at a given time. The uncer-229 191 tainty stated by its manufacturer for individual measurements230 192 is $\pm 8\%$ (± 10 W/m²), both for GHI and DHI at 95% confidence 193 level [22]. Several studies [23-25] have shown that these in-194 struments easily comply with the stated uncertainty for GHI, 195 but their DHI uncertainty can be higher. In a comparison made 196 at NREL in 2009, Myers and Wilcox [23] reported for this in-197 strument uncertainties between 4 to 7% for GHI and 7 to 11% 198 for DHI. Another study [24] compared SPN1 measurements 199 against KZ CM11 pyranometers (one of them equipped with a 200 shadow band) and reported uncertainties of approximately 3% 201 for GHI and 14% for DHI at the 1-minute time scale. More 202 recently, a detailed study by Badosa et al. [25] has compared 203 SPN1 measurements against high quality (sun tracker based) 204 data and found similar uncertainties of approximately 5% for 205 GHI and 12% for DHI. A negative mean bias of approximately 206 5% was also reported for DHI measurements when compared 207 to measurements from a shade-ball assembly. Based on these 208 results, the application of a 1.05 correction factor to the DHI 209 output of this instrument has been recommended [22, 25]. In 210 this work, this correction has been applied to the DHI data from 211 the AR, AZ and TT sites before filtering. 212

Furthermore, we have recently recalled the three SPN1 in-231 213 struments used in this work and calibrated them at our labora-2922 214 tory against two CMP-22 pyranometers (Secondary standards), 223 215 one of them equipped with a shadowring. Additionally, a simul-224 216 taneous measurement of GHI, DHI and DNI based on two new₂₃₅ 217 CMP10 pyranometers and two CHP1 pirheliometers mounted₂₂₆ 218 on a SOLYS2 tracking system where available for consistency₂₃₇ 219 checks. As a result, we have determined that the GHI uncer-228 220 tainty of the three SPN1 instruments is between 3 and 4 % and 220 221 their DHI uncertainty, between 9 and 10 %, without correction₂₄₀ 222 factor. When this factor is included, we have verified that the 223 DHI measurement is essentially unbiased and the uncertainty₂₄₂ 224 in DHI from the SPN1 instruments is between 6 and 7 %, in₂₄₃ 225 agreement with [25]. On this basis, we estimate the SPN1 un-244 226 certainty for GHI at 3 % and for (corrected) DHI at 7 %. 227 245

Based on the uncertainty estimate for each measurement, one₂₄₆

can assign combined uncertainties to the diffuse fraction data for each site, as indicated in the last column of Table 1.



Figure 3: Distributions of (a) hourly k_t values and (b) hourly f_d values, both filtered to F7-level. Data from all sites is shown, since similar distributions are found on a per-site basis.

2.2. Filtering criteria

GHI data separates into its beam (I_{bh}) and diffuse (I_{dh}) components, $I_h = I_{bh} + I_{dh}$. As a first step, the data is normalized in order to remove most trends due to the apparent solar motion. The hourly clearness index, k_t , is defined as $k_t = I_h/I_{0h}$, where I_{0h} is the hourly solar irradiation on a horizontal surface at the top of the atmosphere. The hourly diffuse fraction, f_d , is the ratio $f_d = I_{dh}/I_h$. For cloudy conditions, $k_t \rightarrow 0$ and $f_d \rightarrow 1$. For clear-sky conditions, $k_t \approx 0.80$ and f_d takes low values (≈ 0.10) which depend on the composition of the local atmosphere, as shown in Figs. 2 and 3.

When working with diffuse (or beam) solar irradiation, quality assessment of the data is specially relevant [26, 27]. A filtering procedure is implemented, based on the sequential application of seven filters to the normalized hourly data records, as summarized in Table 2. The process starts with the set (F0)

		A	AZ		1	LU	J	A	R	T	Г	all si	tes
Filter	Conditions	hours	%	hours	%								
F0	$\cos \theta_z \ge 0 \& I_h > 0 \& I_{dh} > 0$	7961		20594		5934		7613		6634		48736	
F1	$\cos \theta_z \ge 0.1219 \ (\alpha_s \ge 7^\circ)$	7062	11.3	18483	10.3	5372	9.5	6974	8.4	5987	9.8	43878	10.0
F2	$I_h \le I_{hc} \left(T_L = 2 \right)$	6863	2.8	18348	0.7	5315	1.1	6909	0.9	5890	1.6	43325	1.3
F3	$k_t > 0.1 \& I_d \ge I_{dc} (T_L = 1.5)$	6796	1.0	18327	0.1	5300	0.3	6909	0.0	5890	0.0	43222	0.2
F4	$I_d \leq (600 \mathrm{W/m^2}) \alpha_s$	6773	0.3	18091	1.3	5142	3.0	6903	0.1	5875	0.3	42784	1.0
F5	$k_t \le 0.10 \& f_d \ge 0.85$	6601	2.5	18045	0.3	5132	0.2	6786	1.7	5853	0.4	42417	0.9
F6	$k_t \le 0.85 \& 0.05 \le f_d \le 1.03$	6559	0.6	17895	1.0	4920	4.1	6558	3.4	5576	4.7	41472	2.2
F7	$\left t\right = \left \hat{f}_d - f_d\right / \sigma < 3$	6491	1.0	17616	1.4	4868	1.1	6486	1.1	5534	0.8	40995	1.2
all	% discarded		18.5		14.5		18.0		14.8		16.6		15.9

Table 2: Sequence of filters applied to the hourly irradiation data from each station. For each site, the number of hours that pass each filter and the percentage of records discarded are indicated. A total of 40995 valid daytime hourly records were used and 15.9 % of the daytime hours were discarded.

of daytime hours with positive global and diffuse hourly irra-271 247 diation records for each site. F1 eliminates hours with low272 248 solar altitude ($\alpha_s < 7^\circ$), for which the measurements become₂₇₃ 249 unreliable. F2 uses the ESRA clear-sky model [28, 29] with274 250 Linke turbidity parameter $T_L = 2$, to provide an upper bound₂₇₅ 251 I_{hc} for GHI. For the hours that pass this filter the hourly clear-276 252 ness index, $k_t = I_h/I_{0h}$, is calculated. Unless for very dark₂₇₇ 253 conditions, diffuse irradiation should be larger than the clear-278 254 sky estimate Idc. Filter F3 uses the same clear-sky model with279 255 $T_L = 1.5$ to apply a lower bound on diffuse irradiation when₂₈₀ 256 $k_t > 0.1$. F4 places an upper bound on diffuse irradiation,₂₈₁ 257 $I_d \leq (600 \,\mathrm{W/m^2}) \,\alpha_s$, (solar altitude expressed in radians), based₂₈₂ 258 in Page's estimate for overcast irradiance [28]. For instance, for₂₈₃ 259 $\alpha_s \approx 80^{\circ}$ this limit is 843 W/m². For the hours that pass this fil-284 260 ter, the diffuse fraction $f_d = I_d/I_h$ is calculated. At overcast sky₂₈₅ 261 conditions the diffuse fraction should be high. F5 removes low₂₈₆ 262 diffuse fractions found at overcast conditions with the require-287 263 ment that if $k_t \le 0.10$ then $f_d \ge 0.85$. On physical grounds, one₂₈₈ 264 would expect $0 \le f_d \le 1$, but these limits are relaxed to account 265 for measurement error. F6 places boundaries on the normalized₂₈₉ 266 data by requiring $0.05 \le f_d \le 1.03$ and $k_t \le 0.85$. The last fil-267 ter, F7, is of a statistical nature and aimed to remove the few²⁹⁰ 268 remaining outliers. A simple polynomial fit (P5) to the F6-level²⁹¹ 269 data is used to compute normalized residuals $t = (\hat{f}_d - f_d)/\sigma$,²⁹² 270 293 where σ is the sample RMSD. Since the residuals are (almost) normally distributed, an hour is considered an outlier (and discarded with 99.7% confidence level) if |t| > 3. Approximately 41000 hourly (k_t , f_d) records result from this procedure, as summarized in Table 2.

The thresholds for all filters have been selected on physical grounds after visual inspection of their effects on the data cloud. It is important to emphasize that the quality of the raw data and the specific choices made in the filtering procedure affect quantitatively the results. The hourly data set filtered to F7-level is shown in Fig. 2 against the background of F0-level data. The distributions for (filtered) k_t and f_d are shown in Fig. 3 (a) and (b) respectively. Note that in (a) the right peak in k_t is associated to clear days while in (b) the right peak in f_d is associated to overcast conditions and the left peak to clear-sky conditions. These distributions are similar to those reported in the literature [6, 30, 31]. Both variables have a bi-modal distribution, although bimodality is weakened at the hourly timescale in k_t .

3. Phenomenological models for hourly diffuse fraction

Phenomenological models attempt to capture the general trend of the diffuse fraction in terms of a set of readily available variables together with basic time and site information. These models are usually adjusted from a limited amount of data for a

few locations. Even though such models are not universal [6], 294 they are often used, at least for engineering purposes [32], over 295 a wide range of locations and atmospheric conditions without 296 information about the associated uncertainties. In this Section, 297 ten well-known hourly diffuse fraction models are introduced. 298 They have been selected with special attention to simplicity and 299 usability and are locally adjusted and evaluated in Section 4. In 300 order to easily refer to them, a short code is assigned to each 301 model as indicated in Table 3. For each model, there is an 302 original version and two locally adjusted versions (per site and 303 global) as described in Section 4. 304

Model	Year	Ref.	Sites	Length	Param.	Pred.	
P5	2015	_	5	40995 h	3	1	-
OH	1977	[33]	1	4 y	2	1	
EKD	1982	[34]	5	5 y	4	1	
RBD	1990	[13]	5	22000 h	9	2	309
SO2	1987	[35]	1	44687 h	6	2	310
BSL	2001	[17, 36]	7	NA	2	1	311
RBL	2010	[14]	7	NA	6	5	312
RA1	2010	[15]	21	$\sim 23 \text{ y}$	4	1	313
RA2s	2010	[15]	21	$\sim 23 \text{ y}$	5	2	314
RA2	2010	[15]	21	$\sim 23 \text{ y}$	7	2	_

Table 3: Models for hourly diffuse fraction considered in this work. NA indicates unknown metadata. "Length" indicates the size of the data set used to train the original model (years or hours); "Param." is the number of adjustable parameters in our implementation and "Pred." is the number of predictor variables.

305 3.1. Simple polynomial model (P5)

A simple polynomial model (P5) for diffuse fraction results from a polynomial function of k_t ,

$$f_d = \begin{cases} 1 & k_t < 0.20 \\ a_0 + a_1 k_t + a_2 k_t^2 + a_3 k_t^3 + a_4 k_t^4 + a_5 k_t^5 & 0.20 \le k_t \le 0.85 \\ c_0 & k_t > 0.85. \end{cases}$$
(2)

subject to continuity constrains for f_d and its derivative f'_d at the endpoints of the central interval. The resulting model has three independent parameters and it serves as a benchmark to



Figure 4: Global polynomial model, Eq. (2) against F7-level data. The coefficients are listed in the rightmost column of Table 4.

evaluate more sophisticated approaches obtained from the literature. This model, adjusted to F6-level data, has been used to compute the residuals for discarding outliers in filter F7, as explained in Section 2.2. The values of the coefficients adjusted to F7-level data, for each site and globally as detailed in Section 4, are listed in Table 4.

P5	AZ	SM	LU	AR	TT	Global
a_0	0.50	0.72	0.80	0.86	1.04	0.77
a_1	5.92	2.80	1.97	0.87	-1.45	2.16
a_2	-22.22	-6.62	-3.93	3.53	13.21	-3.91
a_3	29.51	-4.66	-5.97	-28.43	-43.80	-9.02
a_4	-19.54	14.13	10.96	39.51	48.79	17.00
a_5	6.09	-6.20	-3.56	-16.21	-17.60	-6.79
c_0	0.10	0.09	0.11	0.11	0.12	0.10

Table 4: Parameter sets for model P5, Eq. (2). See Section 4 for details.

3.2. Models OH and EKD

Among the most well-known models are those by Orgill and Hollands [33] (OH) and Erbs et al. [34] (EKD). Both models have been evaluated by several authors previously [6, 7, 37] and can be expressed as

$$f_d = \begin{cases} 1 + a_1 k_t & k_t < k_a \\ b_0 + b_1 k_t + b_2 k_t^2 + b_3 k_t^3 + b_4 k_t^4 & k_a \le k_t \le k_b \\ c_0 & k_t > k_b. \end{cases}$$
(3)

These models differ mainly in their functional form in 316 the central interval, where OH uses a linear expression 317 $(b_2 = b_3 = b_4 = 0)$. Two continuity constrains at $k_a = 0.35$ and 318 $k_b = 0.75$ reduce the number of free parameters in this model 319 to just two. Orgill and Hollands used four years of hourly data 320 from a single site (Toronto, Canada) to obtain the coefficients 321 shown in the first column of Table 5. Emphasizing the local 322 nature of the model, they recommended the use of these param-323 eters for latitudes between 43 °N to 54 °N [33]. 324

OH	Original	AZ	SM	LU	AR	TT	Global
a_1	-0.25	-0.40	-0.29	-0.24	-0.33	-0.19	-0.28
b_0	1.56	1.51	1.60	1.60	1.56	1.63	1.59
b_1	-1.84	-1.86	-2.00	-1.95	-1.93	-1.99	-1.96
c_0	0.18	0.12	0.10	0.14	0.11	0.14	0.12

Table 5: Parameters for model OH, Eq. (3) with $b_2 = b_3 = b_4 = 0$.

For the EKD model uses $k_a = 0.22$ and $k_b = 0.80$ in Eq. (3). 325 The original coefficients for this model, listed in Table 6, were 326 obtained using data from five U.S. sites with latitudes between 327 31 °N and 42 °N with altitudes from 62 m to 1620 m above sea 328 level [34]. Continuity constrains at k_a and k_b and a continu-329 ous derivative at k_b are assumed, so there are four independent 330 parameters. The authors compared this correlation to 3 years 331 of data from Highett, Australia (latitude 38 °S) to evaluate its 332 usefulness in a different climate at a similar latitude. Since 333 then, the EKD model has been used and evaluated world-wide 334 [7, 32, 37, 38] and it has been recommended for universal use 335 in engineering textbooks [32]. Both models (OH and EKD) are 336 shown in the top panels of Fig. 5, in their original and locally 337 adjusted versions, against F7 data. 338

339 3.3. Model RBD

The model by Reindl et al. [13] is an example of a simple, piecewise, multi-predictor model. These authors used 22000

EKD	Original	AZ	SM	LU	AR	TT	Global
a_1	-0.09	-0.24	0.00	-0.06	-0.15	-0.10	-0.09
b_0	0.95	0.70	0.38	0.62	0.68	0.85	0.60
b_1	-0.16	2.63	6.54	3.70	2.91	0.98	3.97
b_2	4.39	-7.38	-21.25	-10.83	-7.75	-0.06	-11.74
b_3	-16.64	1.86	21.37	7.00	1.47	-9.75	7.76
b_4	12.34	2.67	-6.99	-0.30	3.24	8.62	-0.28
<i>c</i> ₀	0.17	0.13	0.09	0.12	0.13	0.13	0.11

Table 6: Parameter sets for model EKD, Eq. (3).



Figure 5: Four single-predictor models. For each case, the original model (dashed line) and the locally adjusted global model (full line) are shown with the data filtered at level F7 in the background.

hours of data from five sites in the U.S. and Europe, with latitudes ranging from 28 °N to 56 °N. An additional set of 3000 hours measured at Oslo, Norway (latitude 60 °N) were used for evaluation purposes. Reindl et al. considered a large set of 28 candidate predictor variables, including those commonly measured at meteorological stations, and used a piecewise linear model to fit the data. They concluded that, on an hourly basis, the best predictor variables were k_i and $\sin \alpha_s$, the sine of the solar altitude. Other relevant predictors might be ambient temperature and relative humidity, but they found this four-predictor model to perform only marginally better than the two-predictor one. Keeping in mind simplicity and usability, we shall consider only the two-predictor version (RBD) defined by,

$$f_d = \begin{cases} a_0 + a_1 k_t + a_2 \sin \alpha_s & f_d \le 1 & k_t < k_a \\ b_0 + b_1 k_t + b_2 \sin \alpha_s & 0.1 \le f_d \le 0.97 & k_a \le k_t \le k_b \\ c_0 + c_1 k_t + c_2 \sin \alpha_s & f_d \ge 0.1 & k_t > k_b, \end{cases}$$
(4)

where $k_a = 0.30$ and $k_b = 0.78$ are fixed. The constrains within each interval are required to avoid unphysical values for possible combinations of the predictors. No continuity constrains are applied. The values of the parameters, as given in Ref. [13], are listed in the first column of Table 7. An evaluation of this model against modern data can be found in Refs. [6, 7, 37].

								. 1
RBD	Original	AZ	SM	LU	AR	TT	Global	
a_0	1.02	0.88	1.00	0.97	0.96	0.96	0.96	
a_1	-0.25	-0.04	-0.09	-0.07	0.12	0.20	-0.01	
a_2	0.01	0.14	0.01	0.02	-0.06	-0.07	0.01	
b_0	1.40	1.39	1.44	1.49	1.42	1.50	1.45	
b_1	-1.75	-1.94	-1.97	-1.96	-2.01	-2.01	-1.98	
b_2	0.18	0.31	0.24	0.22	0.32	0.27	0.26	
c_0	0	0.56	-0.34	-0.15	0.02	-0.38	-0.12	
c_1	0.49	-0.51	0.69	0.43	0.13	0.73	0.38	
<i>c</i> ₂	-0.18	-0.03	-0.12	-0.09	-0.01	-0.10	-0.08	

Table 7: Parameters for RBD model, Eq. (4).

346 3.4. Model SO2

Skartveit and Olseth [35] proposed a piecewise non-linear diffuse fraction model [SO2] which the solar altitude α_s as an additional predictor. In particular, one of the interval boundaries depends on α_s . The diffuse fraction is parametrized as

$$f_d(k_t, \alpha_s) = \begin{cases} 1 & k_t \le k_a & {}_{364} \\ f(k_t, \alpha_s) & k_a \le k_t \le \alpha k_b(\alpha_s) & (5)_{365} \\ f(\alpha k_b, \alpha_s) & k_t \ge \alpha k_b(\alpha_s) & {}_{366} \end{cases}$$

where $\alpha = 1.09$, $k_b(\alpha_s) = r + s \exp(-\alpha_s/\alpha_0)$ and $\alpha_0 = 0.291^{367}$ rad. The non-linear functions are

SO2	Original	AZ	SM	LU	AR	TT	Global
а	0.27	0.06	0.44	0.11	0.10	0.12	0.21
r	0.15	0.01	0.05	0.03	0.04	0.06	0.04
S	0.43	0.64	0.43	0.58	0.42	0.27	0.47
r'	0.87	0.92	0.87	0.91	0.89	0.90	0.90
s'	-0.56	-0.68	-0.53	-0.56	-0.55	-0.43	-0.55
k_a	0.20	0.05	0.24	0.13	0.11	0.14	0.15

Table 8: Parameters for model SO2, Eqs. (5) and (6). The second column is from Ref. [35] and the rest of the columns correspond are fits to the ground data considered in this work.

In Ref. [35], six parameters (a, r, s, r', s', k_a) are obtained from the data. These values, reproduced in the first column of Table 8, are valid for altitudes close to sea level at average snow-free conditions in Norway. This model is continuous at both interval boundaries where it also has an (approximately) continuous partial derivative $\partial f_d / \partial k_t$ at $k_t = \alpha_s k_b$. In spite of its apparent complexity, it has only six adjustable parameters and two predictors. The same authors have proposed a more complex model [16], which includes persistence and ground albedo among other effects, and shall not be considered here.

363 3.5. Logistic models (BSL, RBL)

In 2001 Boland et al. [36] proposed and later evaluated [17] a single predictor model (BSL) using a simple logistic function, with just two parameters derived from data from 8 sites over four continents. More recently [14], Ridley, Boland and Lauret have proposed a multiple-predictor version,

$$f_d = \left[1 + \exp\left(a_0 + a_1 \, k_t + a_2 \, t_s + a_3 \, \alpha_s + a_4 \, K_t + a_5 \,\psi\right)\right]^{-1}.$$
(7)

The two-parameter, single-predictor logistic model (BSL) can be obtained from Eq. (7) by setting $a_2 = a_3 = a_4 = a_5 = 0$ and its original parameters from Ref. [17] are reproduced in Table 9.

The extended model (RBL) considers four predictors aside from k_t : t_s is the apparent solar time (in hours) at the mid-hour point, α_s is the solar altitude angle in degrees, K_t is the daily clearness index and ψ is a persistence parameter defined as the average of the lag and lead hourly clearness index, i.e. for the



Figure 6: Best models (locally adjusted) with multiple predictors. Upper panel: model RBL, Eq. (7); center panel: model SO2, Eq. (5); Bottom panel: model³⁸⁴ RA2, Eq. (8). In the background, the data filtered to F7-level is shown. See³⁸⁵ Section 4 for details on the local adjustment and evaluation of these models.

BSL	Original	AZ	SM	LU	AR	TT	Global
a_0	-5.00	-4.50	-5.03	-4.98	-4.78	-5.23	-4.94
a_1	8.60	8.37	9.26	8.83	8.81	9.18	8.95
RBL							
a_0	-5.38	-5.07	-5.50	-6.02	-5.31	-5.85	-5.60
a_1	6.63	7.29	7.75	7.34	7.94	7.94	7.63
a_2	0.01	-0.03	0.02	0.03	0.00	0.00	0.01
<i>a</i> ₃	-0.01	-0.02	-0.01	-0.01	-0.02	-0.02	-0.01
a_4	1.75	1.12	0.88	1.75	0.68	0.93	1.12
<i>a</i> ₅	1.31	1.92	2.12	1.92	2.08	2.29	2.06

Table 9: Parameters for the logistic models BSL and RBL, Eq. (7). For BSL all parameters a_j with j > 1 are zero. The original values are from Refs. [14, 17].

 j^{th} hour, $\psi(j) = \frac{1}{2} [(k_t(j-1) + k_t(j+1)]]$, unless for sunrise or 373 sunset hours where $\psi(j) = k_t(j \pm 1)$, respectively. Using a per-374 sistence and a daily variable in an hourly model implies that it 375 can not be used in real time or for days with incomplete hourly 376 information. However, it can capture daily trends in the data 377 and offer an improved performance. The original values for 378 these parameters have been obtained in Ref. [14], from data for 379 seven locations worldwide. The coefficients obtained from the 380 data considered in this paper are also shown in Table 9. 381

382 3.6. Double exponential models (RA1, RA2s, RA2)

Ruiz-Arias et al. [15], have proposed the use of a Gompertz (or double exponential) function for diffuse fraction estimation. These functions have previously been used in this context to discard outliers in the (k_t, f_d) plane [20, 27] as they can represent the general trend shown in Fig 2. In this work, we consider one (RA1) and two-predictor (RA2s, RA2) models with the general form

$$f_d = a_0 + a_1 e^{-\exp(a_2 + a_3 k_t + a_4 k_t^2 + a_5 m + a_6 m^2)}$$
(8)

where *m* stands for the height-corrected relative air mass [39] evaluated at the midpoint of the hour. The single-predictor model (RA1) is obtained by setting $a_4 = a_5 = a_6 = 0$ in this expression and a simplified two-predictor model (RA2s) is obtained by setting $a_4 = a_6 = 0$. In Ref. [15] the parameters for each of these models are determined from a large set of high

387

quality data from 21 sites worldwide. These sites range in lati-398
tude from 30 °N to 65 °N and in altitude from sea level to almost
2000 m. The dataset covers a range of climates but all sites are
located in the northern hemisphere.

								402
RA1	Original	AZ	SM	LU	AR	TT	Global	
a_0	0.95	0.95	1.00	0.98	0.95	0.96	0.97	-
a_1	-1.04	-0.97	-1.07	-1.05	-0.92	-0.97	-1.01	
a_2	2.30	2.96	2.82	2.96	3.57	3.46	3.07	
<i>a</i> ₃	-4.70	-6.07	-5.82	-5.75	-7.32	-6.68	-6.17	
RA2s								
a_0	0.98	0.96	1.00	0.97	0.94	0.96	0.97	403
a_1	-1.02	-1.06	-1.17	-1.15	-1.01	-1.08	-1.11	404
a_2	2.88	3.36	3.05	3.32	3.94	3.75	3.38	405
<i>a</i> ₃	-5.59	-5.87	-5.44	-5.56	-6.79	-6.25	-5.84	
a_5	-0.11	-0.15	-0.11	-0.12	-0.17	-0.14	-0.13	406
RA2								407
a_0	0.94	0.97	1.00	0.97	0.94	0.96	0.98	408
a_1	-1.54	-0.94	-1.40	-1.18	-1.11	-1.43	-1.24	409
a_2	2.81	2.72	3.25	3.36	4.30	4.10	3.47	
<i>a</i> ₃	-5.76	-1.01	-6.19	-5.49	-7.58	-7.85	-5.71	410
a_4	2.28	-5.78	1.62	0.10	1.37	2.80	0.32	411
<i>a</i> 5	-0.13	-0.45	-0.19	-0.19	-0.36	-0.22	-0.25	412
<i>a</i> ₆	0.01	0.04	0.01	0.01	0.03	0.02	0.02	_ 413

Table 10: Parameters for the double exponential single-predictor model RA1, based on Eq. (8) with $a_4 = a_5 = a_6 = 0$. The first column is from Ref. [15].

The site-independent set of coefficients recommended for₄₁₇ each model in Ref. [15] are shown in the first column of Table 10 under "Original", together with the corresponding coefficients determined from the data considered in this work.

337 3.7. Evaluation of original models

In Table 3 the ten hourly diffuse fraction models and the length and scope of the data used to train the original version of each model are listed. These original models are evaluated₄₁₈ against the F7-level set of ground measurements (described in₄₁₉ subsection 2.2) and three performance indicators are calculated.₄₂₀ For *n* data points (x_i , y_i) and corresponding estimates \hat{y}_i , the₄₂₁ Mean Bias Deviation (MBD) is defined as

$$MBD = \frac{1}{n} \sum_{i=1}^{n} (\hat{y}_i - y_i), \qquad rMBD = 100 \times \frac{MBD}{\langle y \rangle} \qquad (9)_{424}^{423}$$

where $\langle y \rangle$ is the mean of the observations. It should be noted that this indicator has been defined in both ways (estimate measurement or viceversa) in the literature. Eq. (9) implies that a positive bias is associated to overestimation by a given model, in accordance with standard usage [40].

The Root Mean Square Deviation (RMSD) quantifies the dispersion of the residuals,

RMSD =
$$\left[\frac{1}{n} \sum_{i=1}^{n} (\hat{y}_i - y_i)^2\right]^{\frac{1}{2}}$$
, rRMSD = $100 \times \frac{\text{RMSD}}{\langle y \rangle}$. (10)

For hourly models, the relative forms are scaled using the average of F7-level hourly data, $\langle f_d \rangle = 0.47$. In Section 5, $\langle F_d \rangle = 0.46$ and $\langle \bar{F}_d \rangle = 0.36$ are used to scale the daily and monthly average indicators, respectively. The rRMSD indicator characterizes the uncertainty introduced by the use of a given model to estimate the diffuse component of GHI. Thus, it is relevant to rank the adjusted models according to their capacity to describe the data. The rMBD gives information about systematic tendencies in the models to overestimate or underestimate the data. Combinations of these two indicators, such as Student's *t* parameter [41] or the $\mu_{1-\alpha}$ indicator [42] may be used to rank these original models, with some emphasis on the mean bias indicator. However, since the locally adjusted (essentially unbiased) versions of the models perform significantly better, we make no attempt to rank the original versions.

A Kolmogorov-Smirnov Index (KSI) is defined [4, 40, 43] using the cumulative distribution functions F(Y) and $\hat{F}(Y)$ estimated from the f_d measurements and the corresponding model estimates respectively. KSI quantifies the distance between these distributions,

$$KSI = \int_0^1 D(Y) \, dY$$
, with $D(Y) = \left| \hat{F}(Y) - F(Y) \right|$. (11)

The function *D* helps visualize for which range of f_d the model estimates differ significantly from the data. As an example, in Fig. 8 we show this function for the RBL model. Other models have a similar form for $D(f_d)$ suggesting that their performance under clear sky conditions (low f_d) might be improved, for instance by using an accurate clear-sky model. Thus, KSI gives information about the similarity between the distributions of the

measured and modeled diffuse fractions and discriminates well₄₅₈
 between different models.

These three basic indicators, MBD, RMSD and KSI, are⁴⁶⁰ combined into a single one which takes into account dispersion,⁴⁶¹ absolute bias and likeness of distributions 462

CPI =
$$\frac{1}{3}$$
(|rMBD| + rRMSD + 100 × KSI). (12)⁴⁶³₄₆₄

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This overall indicator is similar to the combined index proposed
 by Gueymard [40].

Meaningful comparisons based on relative indicators calcu-467 429 lated by different authors using different data sets are not al-468 430 ways straightforward. Even when the same data set is used, dif-469 431 ferences in the filtering procedure may affect the performance⁴⁷⁰ 432 indicators obtained for the same model. With this in mind, we⁴⁷¹ 433 note that some of the models considered here have been previ-472 434 ously evaluated elsewhere. Jacovides et al. [10] evaluated the⁴⁷³ 435 single variable models OH, EKD, RBD and BSL (among oth-474 436 ers) using 5-years of hourly data for a semi-rural site in Cyprus.⁴⁷⁵ 437 They reported positive biases between 3-7% and rRMSD in a⁴⁷⁶ 438 narrow range about 30% for all of them. Tapakis [44] used data477 439 from the same site in Cyprus (but for a 13-year period) to eval-478 440 uate models OH, EKD, RBD and RA1 and obtained rMBD in a479 441 narrow range around 5% and rRMSD around 24%. 480 442

These relative indicators are calculated for each original481 443 model, on a per-site basis, in columns 3 to 8 of Table 13.482 444 Note that most models show positive biases (over-estimation483 445 of diffuse fraction) in the range 3-12%. As mentioned, simi-484 446 lar results have been reported previously for Cyprus [10] and⁴⁸⁵ 447 also in a comparison between several separation models with486 448 data from sites closer than 1000 km from our region of interest487 449 [45]. The exceptions to positive biases are the models based on⁴⁸⁸ 450 Gompertz or double exponential functions (RA1, RA2s, RA2),489 451 which show negative biases (under 10%) for all our sites. In490 452 Ref. [15] these models are also reported with mostly negative491 453 biases (between -5% and -12%) when compared to indepen-492 454 dent data from 14 sites in the northern hemisphere. Many of the493 455 models considered here (including RA1 and RA2) have been re-494 456 cently compared (at the 1-minute time scale) against data from495 457

a BSRN site (SMS-São Martino da Serra), located about 500 km to the north of the area of interest of this work [6]. In this comparison, rRMSD between 24-29% and rMBD between 3-7% were found, with the biases from RA1 and RA2 having a different sign than those of the other models. These results are consistent with the left part of Table 13, with rRMSD's in the range 19-28%, depending on model and site. The AZ site, located at the most southern latitude and being the only coastal site in our analysis, has a higher dispersion (rRMSD's) than the rest. This site is representative of a special climatic regime with higher variability, humidity and cloudiness than inland sites.

In terms of global rRMSD, the best original models are RBL and RA2s with 20.7 % and 21.0% respectively, with the last having a lower bias. They are followed by RA2 and SO2 with rRMSD of 21.8% and 22.9%, respectively. Independent work performed in Argentina [11] used 4320 hours of data from one site in a similar climatic region and also found RBL and SO2 (in their original forms) as the two best models in terms of rRMSD. However, this work did not consider Gompertz-based models such as RA2s and RA2. Models RBL and SO2 have also been ranked among the best for Temperate Zones in Ref. [6].

The tendency for overestimation the diffuse fraction found in most models may be due to a clearer atmosphere with fewer aerosols on average, since most sites considered in this work are at semi-rural grasslands with low population density. However, this may also result from imperfect measurements. For instance, no ventilated pyranometers where used, so water droplets on the domes may affect measurements even at high solar altitudes. The use of the isotropic correction for the shadowring diffuse measurements (which affects more than 40% of the data) may produce some underestimation (of the order of 1%) in the diffuse fraction [19]. The exact choices made in the filtering procedure (Table 2) may also affect the mean bias results of the original models. Specifically, if a higher lower limit (i.e. 0.95 instead of 0.85) is chosen for diffuse fraction under cloudy conditions (F5) a slightly smaller bias would be obtained.

The models considered here have been originally adjusted to data mostly from the northern hemisphere and, in some cases,

using small data sets. Since they have significant biases, it
is worth deriving locally adjusted versions of these models,
specifically adapted for this region of the world.

499 **4. Locally adjusted models**

Each model has been trained and evaluated on a per-site ba-500 sis, using a standard cross-validation procedure. At each site, 501 the F7-dataset is randomly divided into a training set (80%) and 502 a testing set (20%) and the optimal set of coefficients for each 503 model is obtained using standard regression techniques. The 504 final coefficients and indicators result from averaging over an 505 ensemble of 1000 elements, each sampling random 80-20 por-506 tions of the datasets. The size of the ensemble has been chosen 507 as to warrant the repeatability of the procedure. 508

	AZ	SM	LU	AR	TT
W _i	0.14	0.32	0.26	0.14	0.14
f_i	0.16	0.43	0.12	0.16	0.13

Table 11: Weights w_i are based on the estimated uncertainty for the data from each site, indicated in Table 1. The second row shows the fraction f_i of data points from each site at F7-level (Table 2).

In addition to the locally adjusted version of each model for 509 each site, global models are constructed from the locally ad-510 justed models using the weighted average of the adjusted co-511 efficients from each site as the global model coefficients. The 512 weights w_i are determined from the estimated uncertainty for 513 each site, σ_i , as indicated in the last column of Table 1. Thus, 514 $w_i = C/\sigma_i^2$ with $C = \sum_i 1/\sigma_i^2$ (the sum runs over all sites). 515 The resulting weight factors, listed in Table 11, give priority to 516 higher quality data from the SM and LU sites. The coefficients 517 for the globally fitted versions of each model are listed in the 518 last column (under Global) of Tables 4 to 10. 519

Performance indicators for these globally adjusted versions are shown in the rightmost columns of Table 13. The per-site indicators are averaged (weighted using the fractions f_i of data points from each site) to obtain the global indicators for each adjusted model, listed under the "All" header in Table 13.

			Adjuste	ed local	models	
Model	Indicator	AZ	SM	LU	AR	TT
	rMBD (%)	1.9	0.0	0.6	2.0	1.5
P5	rRMSD(%)	25.8	19.2	22.5	22.1	22.9
	KSI (×100)	2.3	1.4	1.9	2.2	2.0
	rMBD (%)	0.0	-0.8	-0.1	0.4	0.6
OH	rRMSD(%)	25.8	20.1	22.6	22.2	22.9
	KSI (×100)	2.6	2.3	2.1	2.3	1.9
	rMBD (%)	0.4	0.1	0.3	0.8	0.7
EKD	rRMSD(%)	25.5	19.2	22.5	22.0	22.9
	KSI (×100)	2.2	1.4	1.9	1.9	1.8
	rMBD (%)	0.0	0.0	0.0	0.0	0.0
RBD	rRMSD (%)	23.3	17.3	20.8	18.4	20.5
	<i>KSI</i> (×100)	2.3	1.3	2.0	1.8	1.9
	rMBD (%)	0.9	0.0	0.4	1.5	1.3
SO2	rRMSD (%)	22.9	16.6	20.1	18.2	20.2
	<i>KSI</i> (×100)	1.9	1.2	1.6	1.4	1.6
	rMBD (%)	-0.2	-1.5	-0.6	-0.2	-0.1
BSL	rRMSD (%)	25.6	19.6	22.6	22.1	23.0
	KSI (×100)	2.2	1.6	2.0	1.6	1.6
	rMBD (%)	0.6	-0.8	-0.2	0.7	0.6
RBL	rRMSD(%)	21.2	15.7	17.7	17.4	19.0
	<i>KSI</i> (×100)	1.5	1.2	1.5	1.0	1.3
	rMBD (%)	0.0	0.0	0.0	0.0	0.0
RA1	rRMSD (%)	25.5	19.2	22.4	21.5	22.6
	KSI (×100)	2.5	1.3	2.0	2.1	2.1
	rMBD (%)	-0.1	0.0	0.0	0.0	0.0
RA2s	rRMSD(%)	23.2	16.9	20.3	18.3	20.3
	<i>KSI</i> (×100)	2.3	1.1	1.8	1.8	1.9
	rMBD (%)	0.0	0.0	0.0	0.0	0.0
RA2	<i>rRMSD</i> (%)	22.8	16.6	20.2	17.7	19.9
	KSI (×100)	2.2	1.0	1.8	1.6	1.9

Table 12: Per-site performance indicators for the locally adjusted models as compared to the F7-dataset on a per-site basis. Note that the local Gompertz models (RA1, RA2s and RA2) and the local RBD model are essentially unbiased at all sites. The best local model performance is from RBL at SM, with rRMSD below 16%.

			Original models					Adjusted global models					
Model	Indicator	AZ	SM	LU	AR	TT	All	AZ	SM	LU	AR	TT	All
	rMBD (%)							4.5	2.8	-2.9	3.6	-3.1	1.7
P5	rRMSD (%)							26.1	19.5	22.9	22.2	23.6	21.9
	KSI (×100)							2.3	2.1	2.3	2.5	2.8	2.3
	rMBD (%)	10.7	8.7	4.2	9.6	3.0	7.9	3.8	1.6	-3.7	2.5	-4.0	0.7
OH	rRMSD(%)	28.1	22.9	23.2	24.4	23.4	24.1	26.2	20.3	22.9	22.4	23.4	22.3
	<i>KSI</i> (×100)	5.9	6.0	3.8	5.2	3.6	5.3	2.7	2.9	2.2	2.3	2.2	2.6
	rMBD (%)	10.8	8.8	3.8	9.8	3.3	7.9	4.0	2.2	-3.4	3.0	-3.7	1.2
EKD	rRMSD (%)	28.0	22.2	22.9	24.1	23.1	23.6	25.9	19.6	22.8	22.1	23.5	21.9
	<i>KSI</i> (×100)	5.4	5.0	2.5	4.9	2.3	4.4	2.1	2.4	2.3	2.4	2.7	2.4
	rMBD (%)	11.2	10.6	5.2	11.0	3.8	9.2	3.3	2.3	-3.8	2.6	-4.4	0.9
RBD	rRMSD (%)	26.6	21.9	22.5	22.9	22.7	23.0	23.8	17.9	21.3	18.9	21.4	19.8
	<i>KSI</i> (×100)	5.6	6.0	4.5	6.0	4.2	5.5	2.4	2.7	2.2	2.5	2.5	2.5
	rMBD (%)	13.4	12.2	7.0	12.6	6.2	11.0	3.8	2.3	-4.0	2.5	-3.8	1.0
SO2	rRMSD (%)	26.7	22.1	22.0	22.4	21.9	22.9	23.4	16.9	20.7	18.3	21.1	19.2
	<i>KSI</i> (×100)	6.6	6.6	4.3	6.2	3.7	5.8	1.9	2.2	2.5	1.5	2.3	2.1
	rMBD (%)	11.5	9.3	4.7	10.2	3.6	8.5	3.0	1.1	-4.3	1.9	-4.7	0.1
BSL	<i>rRMSD</i> (%)	28.6	23.3	23.3	24.9	23.5	24.4	25.9	19.9	23.0	22.2	23.6	22.1
	<i>KSI</i> (×100)	6.0	6.2	3.7	5.2	3.3	5.3	2.2	2.7	2.4	1.8	2.6	2.4
	rMBD (%)	10.3	8.7	2.5	8.9	3.0	7.6	3.8	2.7	-4.7	2.4	-3.4	1.3
RBL	<i>rRMSD</i> (%)	25.0	19.8	18.7	20.9	20.4	20.7	22.2	16.3	18.8	17.8	19.5	18.1
	<i>KSI</i> (×100)	5.3	5.6	3.2	4.6	3.2	4.9	2.0	2.4	2.6	1.4	1.9	2.1
	rMBD (%)	0.0	-1.1	-5.8	-0.4	-7.6	-2.2	3.4	1.5	-3.8	2.4	-4.3	0.5
RA1	rRMSD (%)	26.0	21.3	24.3	22.8	25.6	23.2	25.8	19.6	22.8	21.8	23.3	21.8
	KSI (×100)	4.4	5.2	5.0	4.4	5.6	5.0	2.1	2.3	2.5	2.1	2.8	2.3
RA2s	rMBD (%)	-1.2	-1.8	-7.5	-1.5	-8.5	-3.2	3.8	1.7	-4.0	2.2	-4.1	0.7
	rRMSD (%)	23.9	18.5	23.3	20.2	24.5	21.0	23.7	17.4	20.8	18.5	21.0	19.5
	KSI (×100)	3.7	4.3	5.3	3.6	5.6	4.4	2.3	2.2	2.3	1.7	2.4	2.2
	rMBD (%)	-2.5	-3.1	-8.4	-2.7	-10.1	-4.5	0.9	-0.9	-7.4	-0.6	-7.0	-2.2
RA2	<i>rRMSD</i> (%)	24.6	19.7	23.8	20.9	24.8	21.8	23.5	17.0	21.9	18.3	21.8	19.5
	<i>KSI</i> (×100)	4.0	4.5	5.3	3.7	5.9	4.6	2.3	2.1	2.3	2.5	2.8	2.3

Table 13: Performance indicators for the original and globally adjusted versions of the hourly models. The indicators listed under the columns labelled "All", are the average of the per-site indicadors, weighted by the fraction of data points at each site (see Table 2).



Figure 7: Scatter plot for the diffuse fraction from the RBL model: (a) model₅₅₉ with original coefficients and (b) global model with adjusted coefficients. A line with slope 1 is drawn to guide the eye.

As expected, the adjusted versions have significantly lower563 525 biases than the original ones and the global adjusted versions564 526 have biases within ±2%. Original models have rRMSD indica-565 527 tors in the range 18-28% and locally fitted models in the range566 528 16-26%, so the improvement in dispersion is not as significant⁵⁶⁷ 529 as with bias. In Table 14, all models are listed as ordered by 568 530 increasing CPI, defined in eq. (12). Ties in CPI are resolved₅₆₉ 531 by KSI or rRMSD (both yield the same result). The procedure₅₇₀ 532 is fairly robust: the first two models have the lowest KSI and₅₇₁ 533 rRMSD and any single variable model performs worst than any₅₇₂ 534 multi-variable model also in terms of KSI or rRMSD. 535 573

The best performing global model for this region is clearly⁵⁷⁴ RBL with CPI of 7.2 and rRMSD of 18.1%. It is followed⁵⁷⁵

⁵³⁸ by SO2 (19.2%) and RA2s tied with CPI=7.4 (RA2s has higher
⁵³⁹ rRMSD 19.5% and lower bias than SO2), however RA2s is sim⁵⁴⁰ pler to implement. RA2 and RBD follow tied with CPI=7.8, the
⁵⁴¹ first has lower rRMSD and KSI.

as noted, all multiple-predictor models perform better than 542 any single-predictor model based on k_t alone, so it is worth in-543 cluding air mass m and other additional variables as predictors 544 for describing diffuse fraction in this region. All local single-545 predictor models have rRMSD indicators around 22%, imply-546 ing that the details of these models are not very relevant, as long 547 as they are limited to k_t as the only predictor. In order to empha-548 size this point, we have introduced a simple fifth degree polyno-549 mial model with natural constrains (P5) which is ranked worst 550 in terms of CPI but in terms of rRMSD and rMBD actually per-551 forms almost as well as the best adjusted single-predictor model 552 (RA1). 553

Scatter plots for the original and globally adjusted versions of the RBL model are shown in Fig. 7 and the difference function D used to calculate the distance between the distributions from the data and the RBL model estimates (for the original and globally adjusted versions) is shown in Fig. 8. The effect of the local adjustment is seen in the reduction of the area under the difference function. Similar results are obtained for other models. The peak at low f_d values (clear sky conditions) suggests that some improvement in the model's performance may result if the low-end f_d estimation is done by using a locally tuned clear-sky model for diffuse radiation, such as [29]. Fig. 8 also shows a peak in D when $f_d \rightarrow 1$, so there is room for improvement at overcast conditions too. This potential for improvement should be considered when addressing the subject of improved physical models for diffuse fraction.

The RBL model, Eq. (7), stands out in its use of the daily clearness index (K_T) and a persistence parameter (which depends on the previous and the next hour) as predictors. This particular parametrization is probably related to its good performance. However, the use of a daily variable makes it inadequate for real-time (on-the-fly) estimation of hourly diffuse irradiation or for short-term forecasting applications. The second-best

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	rMB	D (%)	rRMS	SD (%)	KSI (CPI	Rank	
Model	Original	Adjusted	Original	Adjusted	Original	Adjusted	Adj	usted
RBL	7.6	1.3	20.7	18.1	4.9	2.1	7.2	1
SO2	11.0	1.0	22.9	19.2	5.8	2.1	7.4	2
RA2s	-3.2	0.7	21.0	19.5	4.4	2.2	7.4	2
RA2	-4.5	-2.2	21.8	19.5	4.6	2.3	7.8	3
RBD	9.2	0.9	23.0	19.8	5.5	2.5	7.8	3
RA1	-2.2	0.5	23.2	21.8	5.0	2.3	8.2	4
BSL	8.5	0.1	24.4	22.1	5.3	2.4	8.2	4
EKD	7.9	1.2	23.6	21.9	4.4	2.4	8.5	5
OH	7.9	0.7	24.1	22.3	5.3	2.6	8.5	5
P5	-	1.7	-	21.9	-	2.3	8.7	6

Table 14: Overall performance indicators for the ten hourly models considered in this work in their original and adjusted versions. The "Rank" column orders the adjusted models using a combined performance indicator defined in Eq. (12). The horizontal line separates multiple-predictor models from single-predictor models.



Figure 8: Difference between cumulative distribution functions, D, from Eq. (11), for the original (dashed line) and the adjusted global (full line) versions of the RBL model.

⁵⁷⁶ models are SO2 and RA2s with similar indicators. The adjusted ⁵⁷⁷ RA2s model is almost unbiased and has a simpler parametriza-⁵⁷⁸ tion than SO2. Thus, for the average user, the locally fitted⁵⁹¹ ⁵⁷⁹ RA2s model, Eq. (8), which can estimate hourly f_d with an⁵⁹² ⁵⁸⁰ uncertainty under 20%, may represent the best compromise be-⁵⁹³

⁵⁸¹ tween performance and simplicity.

2 5. Daily and monthly-mean diffuse fraction

In real applications, daily data for GHI may be the only information available near the required location. In order to estimate the daily solar resource on an inclined surface, a separation into daily diffuse and direct irradiation is previously required. In this Section, two models which have been widely used to obtain this separation are evaluated, in their original and locally adjusted versions. The case of monthly means of daily irradiation is also discussed.

In terms of the daily global and diffuse horizontal irradiation (H_h, H_{dh}) and the daily horizontal irradiation at the top of the atmosphere (H_{0h}) , the daily clearness index and the daily diffuse fraction are defined as,

$$K_T = H_h/H_{0h}$$
 and $F_d = H_{dh}/H_h$. (13)

The monthly-averaged versions of these quantities are

$$\overline{K}_T \equiv \overline{H}_h / \overline{H}_{0h}$$
 and $\overline{F}_d \equiv \overline{H}_{dh} / \overline{H}_h$, (14)

where the averages are over daily data within each month.

For daily and monthly averaged daily data, the specific sitedependence is weaker than for hourly data. For simplicity



Figure 9: Top: CPR-d model for daily data, Eq. (15), (original and adjusted) against the bachground of the daily data. Middle: EKD-d model for daily data (original and adjusted), Eq. (16), with $\omega_s \leq 81.4^{\circ}$ (winter). Bottom: EKD-d model for daily data (original and adjusted) with $\omega_s > 81.4^{\circ}$.

and brevity, daily data from all sites is aggregated and a single locally adjusted version of each model is considered. Our
objective is to assess the typical uncertainty associated to the
diffuse-direct separation procedure at the daily and monthly
time scales.

599 5.1. Data base

Daily data is obtained from hourly data as follows. For a 600 given day, the subset of hours that pass F0, F1, F2 and F6 are 601 considered. For days with complete hours, hourly data is ac-602 cumulated to generate daily irradiation in the usual form, i.e. 603 $H_h = \sum_j I_h(j)$ and $H_{dh} = \sum_j I_{dh}(j)$, where j is an hour index 604 and the sums run over all daylight hours. Days with one or 605 more missing hours are discarded. If a given month has at least 606 20 days with daily data, the monthly averages, \bar{K}_T and \bar{F}_d , are 607 computed otherwise the month is discarded. The results of this 608 selection process are summarized in Table 15. 609

Site Code	Valid days	Valid months
AZ	549	16
SM	1538	53
LU	375	12
AR	506	18
TT	417	13
All sites	3385	112

Table 15: Summary of the filtered daily data for each station and valid months for calculating monthly-averages.

610 5.2. Models

Two frequently used models are considered in their daily and monthly averaged versions. The daily model by Collares-Pereira et al. [46] (CPR-d), is defined in four intervals in K_T ,

$$F_{d} = \begin{cases} 1 & K_{T} \le 0.17 \\ A_{0} + A_{1} K_{T} + A_{2} K_{T}^{2} + A_{3} K_{T}^{3} + A_{4} K_{T}^{4} & 0.17 < K_{t} \le 0.75 \\ B_{0} + B_{1} K_{T} & 0.75 < K_{T} < 0.80 \\ C_{0} & K_{T} \ge 0.80. \end{cases}$$
(15)

where the original and locally adjusted values of the coefficients are listed in Table 16. The daily model from Erbs et al. [34] (EKD-d) is also considered,

$$F_d = \begin{cases} 1 + A_1 K_T + A_2 K_T^2 + A_3 K_T^3 + A_4 K_T^4 & K_T < 0.715 \\ B_0 & K_T \ge 0.715. \end{cases}$$
(16)

with coefficients listed in Table 16. Note that it includes a seasonal dependence through ω_s , the sunset hour angle: its coefficients have different values for ω_s below (i.e. winter) or above (rest of the year) a threshold of 81.4°.

The monthly-average model proposed by Collares-Pereira [46] (CPR-m) uses the monthly averaged hour angle $\bar{\omega}_s$ (in rads) to introduce seasonal dependence,

$$\overline{F}_d = A - B \cos(A_2 + A_3 \overline{K}_T) \tag{17}$$

where $A = A_0 + A_1 \left(\bar{\omega}_s - \frac{\pi}{2}\right)$ and $B = B_0 + B_1 \left(\bar{\omega}_s - \frac{\pi}{2}\right)$. The monthy mean sunset angle, $\bar{\omega}_s$, can be approximated by its value for the typical day of each month [47] with negligible error. The original and adjusted values for these coefficients are listed in Table 16.

Finally, the model for monthly-averaged diffuse fraction (EKD-m) by Erbs et al. [34] is also considered. It is defined by

$$\overline{F}_d = A_0 + A_1 \overline{K_T} + A_2 \overline{K_T}^2 + A_3 \overline{K_T}^3,$$
(18)

with two sets of coefficients according to the value of $\bar{\omega}_s$. The 620 monthly averaged clearness index is restricted to the interval 621 $0.3 \le \overline{K_T} \le 0.8$ and the coefficients are listed in Table 16. The 622 locally adjusted EKD-m model for $\omega_s \leq 81.4$ has some insta-623 bilities. On the other hand, inspection of Fig. 10 shows that 624 the F_d data considered in this work has only weak seasonal de-625 pendence. Ignoring this dependence results in a stable locally 626 adjusted model with similar performance indicators as those 627 obtained by preserving the ω_s dependence. Thus, the locally 628 adjusted EKD-m version does not include the ω_s dependence 629 and a single set of local coefficients are listed in the last row of 630 Table 16. 631

632 5.3. Evaluation

After a cross-validation procedure similar to the one used for the hourly models, performance indicators are obtained for the



Figure 10: CPR model for monthly-averaged data, Eq. (17). Monthly averaged values are indicated within brackets $\langle \cdot \rangle$. EKD model for monthlyaveraged data, Eq. (18). Data for $\bar{\omega}_s \leq 81.4^{\circ}$ is shown with blue circles and with $\bar{\omega}_s > 81.4^{\circ}$ with yellow circles. For each case, the original EKD monthly model, Eq. (16), is shown with dotted lines. The locally fitted model (with no dependence with ω_s) is shown with a full line. Monthly averaged values are indicated within brackets.

		Original models							
Model	restriction	A_0	A_1	A_2	A_3	A_4	B_0	B_1	C_0
CPR-d		1.19	-2.27	9.47	-21.87	14.65	0.63	-0.54	0.20
EKD-d	$\omega_s \leq 81.4^\circ$	1	-0.27	2.45	-11.95	9.39	0.14		
	$\omega_s > 81.4^\circ$	1	0.28	-2.56	0.85	0	0.18		
CPR-m		0.78	0.35	2.00	-1.80	*	0.51	0.26	*
FVD	$\omega_s \leq 81.4^\circ$	1.39	-3.56	4.19	-2.14				
EKD-III	$\omega_s > 81.4^\circ$	1.31	-3.02	3.43	-1.82				
		Adjusted models							
CPR-d		1.49	-5.05	16.89	-29.15	16.52	0.79	-0.89	0.08
EKD-d	$\omega_s \le 81.4^\circ$	1	0	-0.46	-4.50	3.89	0.13		
EKD-d	$\omega_s > 81.4^\circ$	1	0	-1.88	0.34	0	0.15		
CPR-m		6.21	0.56	0.93	-0.67	*	5.95	0.57	*
EKD-m		1.58	-3.67	2.68	-0.19				

Table 16: Original and locally adjusted parameters for daily and monthly-average models for diffuse fraction.

adjusted models. The performance indicators for the original₆₅₅ 635 and adjusted daily models are shown in Table 17. As expected,656 636 large bias indicators (between 5% and 10%) are obtained and at657 637 the daily timescale the original models also overestimate daily658 638 diffuse fraction in the region of interest. For both models, the659 639 local fit reduces the mean bias (below 1%) and leads to lower₆₆₀ 640 F_d estimates, as shown in Fig. 9. The locally fitted versions₆₆₁ 641 of both daily models perform similarly, with a small edge for₆₆₂ 642 EKD-d, estimating daily diffuse irradiance with rRMSD under663 643 20% and the same KSI. 644

The monthly-averaged data, together with the estimates from⁶⁶⁴ 645 the two models considered, are shown in Fig. 10. Due to the 646 small size of the monthly dataset, all of it was used to fit and 647 222 evaluate the monthly mean models. Even though this leads 648 to artificially lower performance indicators, they are useful to 649 668 compare the adjusted model to its original version and to com-650 660 pare monthly models between themselves. 651 670

Again, both original models tend to overestimate monthly-671 mean diffuse irradiation, as they do at the daily and hourly time672 scales and have large rRMSD (for averaged quantities) between673 16% and 20%, with EKD-m outperforming CPR-m. The locally adjusted versions of both models are essentially unbiased and their rRMSD are significantly lower (under 13 %) as indicated in Table 17. As in the daily case, both adjusted models perform equally well.

It is worth using local adjusted models at the daily and monthly-average scales in order to reduce bias. The minimum uncertainties introduced when using the unbiased versions are 20% (daily) and 13% (monthly), respectively.

6. Conclusions

The uncertainty introduced by phenomenological models for diffuse fraction separation has been well characterized for a temperate region located in the southern part of South America. The daytime hourly data was quality assessed and almost 41000 hours of valid data from five sites (most of them in semirural areas) are the basis for this work. Ten models for hourly diffuse fraction have been implemented and evaluated in their original and locally adjusted forms. Half of the models considered use the clearness index as their single variable and the

	rMBD (%)		rRMS	D (%)	KSI (2	x100)	CPI		
Model	Orig.	Adj.	Orig.	Adj.	Orig.	Adj.	Orig.	Adj.	
EKD (d)	5.5	0.4	20.6	19.7	2.7	1.0	9.6	7.0	
CPR (d)	10.2	0.7	22.8	19.9	4.9	1.0	12.6	7.2	
EKD (m)	7.0	0.0	16.1	12.8	3.4	1.1	8.8	4.6	
CPR (m)	9.8	0.0	20.2	12.8	4.8	1.1	11.6	4.6	

Table 17: Statistical indicators for the daily and monthly mean diffuse fraction models considered in this work. For each model, the indicators with the original and the locally fitted coefficients are shown. The average daily diffuse fraction is $F_d = 0.46$ and the monthly average diffuse fraction is $\overline{F}_d = 0.36$.

other half includes other variables as predictors.

The five multi-variable models outperform, in terms of dis-703 675 persion, any of the single variable models considered, so the⁷⁰⁴ 676 best original models are the multi-variable models with uncer-705 677 tainties of at least 21%. Most original models over-estimate⁷⁰⁶ 678 the diffuse fraction with biases in the range 3-12%, depending⁷⁰⁷ 679 on the site. Gompertz based (double exponential) models are⁷⁰⁸ 680 the exception and have small negative biases. These results are⁷⁰⁹ 681 dependent on the quality of the experimental data. This may⁷¹⁰ 682 be a true effect due to a clearer atmosphere, which is plausi-711 683 ble given the geographical characteristics and the relatively low⁷¹² 684 industrialization and human density of the area under consid-713 685 eration. But the possibility that it is due to some residual bias⁷¹⁴ 686 present in the data after the filtering process cannot be ruled out₇₁₅ 687 at present. Further work is required, based on higher quality₇₁₆ 688 data for the area, before this overestimation can be confirmed. 717 689

A locally adjusted and a global version of each model where⁷¹⁸ 690 obtained and evaluated per-site using cross validation proce-719 691 dures. Our results clearly show that multiple-predictor mod-720 692 els perform consistently better that any single-predictor ones,721 693 both in their original and local versions. The adjusted models722 694 do not show the overestimation tendency present in the original 695 models. Mean biases are within $\pm 5\%$ and within $\pm 2\%$ for the 696 global versions. The adjusted versions span a range of $rRMSD_{725}$ 697 between 16% and 26%, depending on site and model. 698

Using a combined performance indicator which takes into₇₂₇ account bias, dispersion and similarity between the data and₇₂₈ the modeled distributions, the ten adjusted hourly models have₇₂₉ been ranked according to their overall performance in the region under consideration. The best of them, RBL [14], can estimate hourly diffuse fraction in the region of interest with a typical uncertainty of 18% and 1% bias. However, this model uses daily irradiation as an input and can't be used for real-time (on demand) separation or for predictive purposes. On a second level, are the SO2 and RA2s models [15, 16] with typical uncertainty under 20% and negligible bias. These models do not share the limitation of the RBL in regard to real-time use. The adjusted RA2s (double exponential) model has a simpler parametrization than the SO2 and, for the average user, RA2s may represent the best compromise between performance and simplicity.

At the daily and monthly mean timescale, two models (CP and EKD) were evaluated before and after adjusting them to to local data. In their original forms both tend to overestimate diffuse irradiation. In their adjusted versions, both daily models are essentially unbiased and perform similarly with typical uncertainty under 20%. In the monthly average case, both adjusted models are indistinguishable, with typical dispersion of about 13%.

In engineering applications, the overall uncertainty introduced by the diffuse radiation estimation should be carefully included in the calculations. The rather high biases found in some original models imply that caution is required before using phenomenological diffuse fraction models outside the regions for which their coefficients where estimated, even at similar latitudes or at a-priori similar climates, since average atmospheric composition may be different due to natural or human-related⁷⁴⁵
causes. Ideally, a local assessment of a proposed model against
good quality local diffuse irradiation data should be considered.⁷⁴⁶

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References

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749

750

- [1] C. Gueymard, Solar Energy 82 (2007) 272.
- [2] G. Abal et al., Anales de la IV Conferencia Latinoamericana de Energía Solar (IV ISES-CLA), pp. 1–12, Universidad Nacional del Cusco, Perú, 2010.
- [3] R. Alonso Suárez et al., Annals of the Solar World Congress (SWC 2011), Kassel, Germany, 2011.
- [4] R. Alonso Suárez et al., Solar Energy 86 (2012) 3205.
- [5] R. Alonso Suárez, M. D'Angelo and G. Abal, Proceedings IV Congresso Brasileiro de Energia Solar e V Conferencia Latino-Americana da ISES, Sao Paulo, Brazil, 2012.
- [6] C. Gueymard and J. Ruiz-Arias, Solar Energy 128 (2016) 1.
- [7] S. Dervishi and A. Mahdavi, Solar Energy 86 (2012) 1796.
- [8] H. Li et al., Renewable Energy 36 (2011) 1944.
- [9] R. Tapakis, S. Michaelides and A. Charalambides, Solar Energy (2014).
- [10] C. Jacovides et al., Renewable Energy 31 (2006) 2492.
- [11] C. Raichijk and F. Taddei, Avances en Energías Renovables y Medio Ambiente - ASADES 16 (2012) 11.23.
- [12] C. Gueymard, Solar Energy 83 (2009) 432.
- [13] D.T. Reindl, W.A. Beckman and J.A. Duffle, Solar Energy 45 (1990) 1.
- [14] B. Ridley, J. Boland and P. Lauret, Renewable Energy 35 (2010) 478.
- [15] J. Ruiz-Arias et al., Energy Conversion and Management 51 (2010) 881.
- 767 [16] A. Skartveit, J.A. Olseth and M.E. Tuft, Solar Energy 63 (1998) 173.
- ⁷⁶⁸ [17] J. Boland, B. Ridley and B. Brown, Renewable Energy 33 (2008) 575.
- [18] A. Drummond, Arch. für Meteorologie, Geophysik und Bioklimatologie,
 Ser. B 7 (1956) 413.
- [19] G. Sánchez et al., JOURNAL OF GEOPHYSICAL RESEARCH 117(2012) D09206.
- [20] C. Raichijk, Avances en Energías Renovables y Medio Ambiente 16(2012) 11.17.
- ISO Technical Committee ISO/TC 180, Solar Energy Sub-committee
 SC1: Climate, Measurement and Data, International Organization for
 Standarization (ISO) preprint ISO 9847:1992(E) (1992).
- [22] J. Wood, Delta-T Devices preprint Rev. 3 (2015), also: private commu nication by Stephen Williams from Delta-T Devices.
- [23] D. Myers and S. Wilcox, National Renewable Energy Laboratory preprint
 NREL/CP-550-45374 (2009).
- [24] B. Psiloglou, S. Lykoudis and D. Kouvas, Advances in Meteorology,
 Climatology and Atmospheric Physics (Springer-Verlag, 2012) chap. ;
 Performance Assessment of an integrated sensor for simultaneous mea surements of global and diffuse radiation components at Athens area, pp.
 259–264.
- 787 [25] J. Badosa et al., Atmos. Meas. Tech. 7 (2014) 4267.
- 788 [26] M. Journée and C. Bertrand, Solar Energy 85 (2011) 72.
- 789 [27] S. Younes, R. Claywell and T. Muneer, Energy 30 (2005) 1533.
- [28] T. Muneer, Solar radiation and daylight models, 2nd ed. ed. (Elsevier-Butterworth-Heineman, Oxford, 2004).

⁷⁹² [29] C. Rigollier, O. Bauer and L. Wald, Solar Energy 68 (2000) 33.

821 Appendix A. Glossary

[30] J. Tovar-Pescador, Modelling the Statistical Properties of Solar Radiation
 and Proposal of a Technique Based on Boltzmann Statistics, Modeling
 Solar Radiation at the Earth's Surface: Recent Advances, edited by V.

⁷⁹⁶ Badescu, chap. 3, pp. 55–91, Springer, 2008.

- [31] A. Ianetz and A. Kudish, A Method for Determining the Solar Global
 and Defining the Diffuse and Beam Irradiation on a Clear Day, Modeling
 Solar Radiation at the Earth's Surface: Recent Advances, edited by V.
- Badescu, chap. 4, pp. 93–112, Springer, 2008.
- [32] J. Duffie and W. Beckman, Solar Engineering of Thermal Processes,
 Third ed. (Wiley and Sons, Hoboken, New Jersey, 2006).
- ⁸⁰³ [33] J. Orgill and G. Hollands, Solar Energy 19 (1977) 357.
- ⁸⁰⁴ [34] D. Erbs, S. Klein and J. Duffie, Solar Energy 28 (1982) 293.
- 805 [35] A. Skartveit and J. Olseth, Solar Energy 38 (1987) 271.
- [36] J. Boland, L. Scott and M. Luther, Environmetrics 12 (2001) 575.
- 807 [37] C. Jacovides et al., Renewable Energy 35 (2010) 1820.
- ⁸⁰⁸ [38] R. Perez et al., ASHRAE Transactions 98 (1992) 354.
- ⁸⁰⁹ [39] F. Kasten and A. Young, Applied Optics 28 (1989) 4735.
- [40] C. Gueymard, Renewable and Sustainable Energy Reviews 39 (2014)
 1024.
- 812 [41] R. Stone, Solar Energy 51 (1993) 289.
- [42] M. de Simón-Martín, M. Diez-Mediavilla and C. Alonso-Tristán, Solar
 Energy (2016), doi 10.1016/j.solener.2016.09.026.
- 815 [43] B. Espinar et al., Solar Energy 83 (2009) 118.
- 816 [44] R. Tapakis, A. Michaelides and A. Charalambides, Solar Energy (2015).
- ⁸¹⁷ [45] C. Raichijk and A. Fasulo, Avances en Energías Renovables y Medio
- 818 Ambiente ASADES 13 (2009) 11.17.
- [46] M. Collares-Pereira and A. Rabl, Solar Energy 22 (1979) 155.
- ⁸²⁰ [47] S. Klein, Solar Energy 19 (1977) 325.

Symbol	Name	Unit
GHI	global horizontal irradiance	Wm^{-2}
DHI	diffuse horizontal irradiance	Wm^{-2}
DNI	beam or direct normal irradiance	Wm^{-2}
I_h	global horizontal hourly irradiation	Whm ⁻²
I_{dh}	diffuse horizontal hourly irradiation	Whm^{-2}
I_{bh}	beam horizontal hourly irradiation	Whm ⁻²
f_d	hourly diffuse fraction = I_{dh}/I_h	
I_{0h}	extraterrestrial hourly horizontal irradiation	Whm^{-2}
k_t	hourly clearness index = I_h/I_{0h}	
θ_z	solar zenith angle	rad
α_s	solar altitude angle	rad
δ	solar declination angle	rad
ϕ	latitude	rad
I_{dc}	clear-sky diffuse hourly horizontal irradiation	Whm ⁻²
I_{bc}	clear-sky beam hourly irradiation	Whm ⁻²
I_c	clear-sky global horizontal irradiation	Whm ⁻²
I_{sc}	hourly solar constant = 1367	Whm ⁻²
ϵ	eccentricity of the earth's orbit	
т	air mass	
T_L	Linke Turbidity at $m = 2$	
δ_R	Rayleigh optical thickness	
T_{rd}	diffuse transmittance function	
F_{da}	diffuse angular function	
H_{0h}	extraterrestrial daily irradiation = $\sum_{day} I_{0h}$	MJm^{-2}
H_h	global daily horizontal irradiation = $\sum_{day} I_h$	MJm^{-2}
K_T	daily clearness index = H_h/H_{0h}	
ω_s	sunset hour angle	rad
H_{dh}	diffuse daily horizontal irradiation = $\sum_{day} I_{dh}$	MJm^{-2}
F_d	daily diffuse fraction	
\overline{H}_{0h}	monthly mean extraterrestrial daily irradiation	MJm^{-2}
\overline{H}_h	monthly mean global daily horizontal irradiation	MJm^{-2}
\overline{K}_T	monthly mean clearness index = $\overline{H}_h/\overline{H}_{0h}$	
\overline{H}_{dh}	monthly mean diffuse daily horizontal irradiation	MJm ⁻²
\overline{F}_d	monthly mean diffuse fraction = $\overline{H}_{dh}/\overline{H}_h$	